

Title: Thoughts on vacuum growth

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Abstract:





- Vacuum effects may violate energy conservation
  - mode creation (Danielsson, ...)
  - causal fluctuations (Sorkin, ...)

◦ Bianchi identity  $G_{ab} = 8\pi G T_{ab}$

$$\nabla^a G_{ab} \equiv 0 \Rightarrow \nabla^a T_{ab} = 0$$

must relax Einstein eq'n.

◦ In full symmetry  $G_{tt} = 8\pi G T_{tt} \rightarrow H^2 = \frac{8\pi G}{3} \rho$ .

$$G_{rr} = 8\pi G T_{rr} \rightarrow \ddot{a} = \dots$$

$$\nabla^a T_{at} = 0 \rightarrow \dot{\rho} + 3\frac{\dot{a}}{a}(\rho + p) = 0$$

So  $G_{rr}$  can be replaced by  $\nabla^a T_{at} = 0$ .

If  $\nabla^a T_{at} = q \neq 0$  can use this instead...

- Covariant formulation of this: drop trace (Danielsson) of Einstein eq'n, i.e. impose only

$$G_{ab} = 8\pi G T_{ab} + \Phi g_{ab}, \quad \Phi \text{ arbitrary function}$$

follows from "unimodular gravity" ( $\det g = \text{fixed}$ )

or from derivation of Einstein eq'n as eq'n of state.

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- But violation of energy conservation cannot be arbitrary:

$$\delta \pi G \nabla^a T_{ab} + \nabla_b \Phi = 0.$$

i.e.  $\nabla^a T_{ab}$  is gradient of a scalar (\*)

→ Automatic for RW symmetry

→ fails already for Bianchi models (anisotropic) let alone generally.

- Is there a way to guarantee (\*) ?

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